

# Future beam options for long baseline neutrino experiments



The 4th International Workshop on Nuclear and Particle Physics at J-PARC

# NP08

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# Outline

1. On the road to the study of the last piece of the neutrino mixing matrix: experimental status
2. Expected performance of T2K phase 1
3. Landscape of the neutrino program  
at the end of T2K phase 1
4. Which strategy for T2K phase 2  
to exploit fully the result of T2K phase 1
5. Conclusions

# Caveat

I am considering myself as a candid in this workshop,  
Apologies if raising obvious questions and any possible misunderstanding

# The missing piece of the neutrino mixing matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\theta_{\text{atm}}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\theta_{13}, \delta} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\theta_{\text{sol}}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Atmospheric sector  
Ongoing studies:

- MINOS :  $\sin^2 2\theta_{23}$ ,  $\Delta m_{23}$
- OPERA :  $\tau$  appearance

Where is the Graal :

- Mixing angle
- CP phase
- mass hierarchy

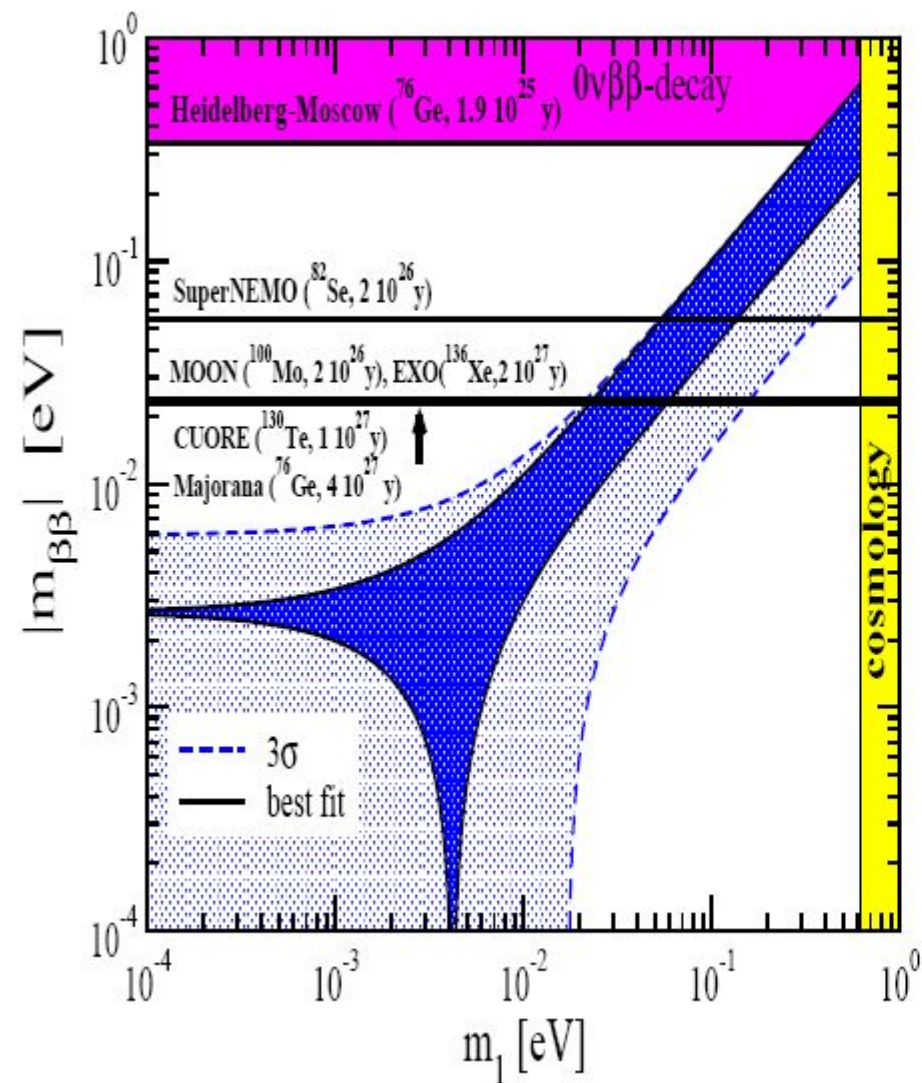
Solar sector  
Ongoing studies:

- KamLAND
- BOREXINO

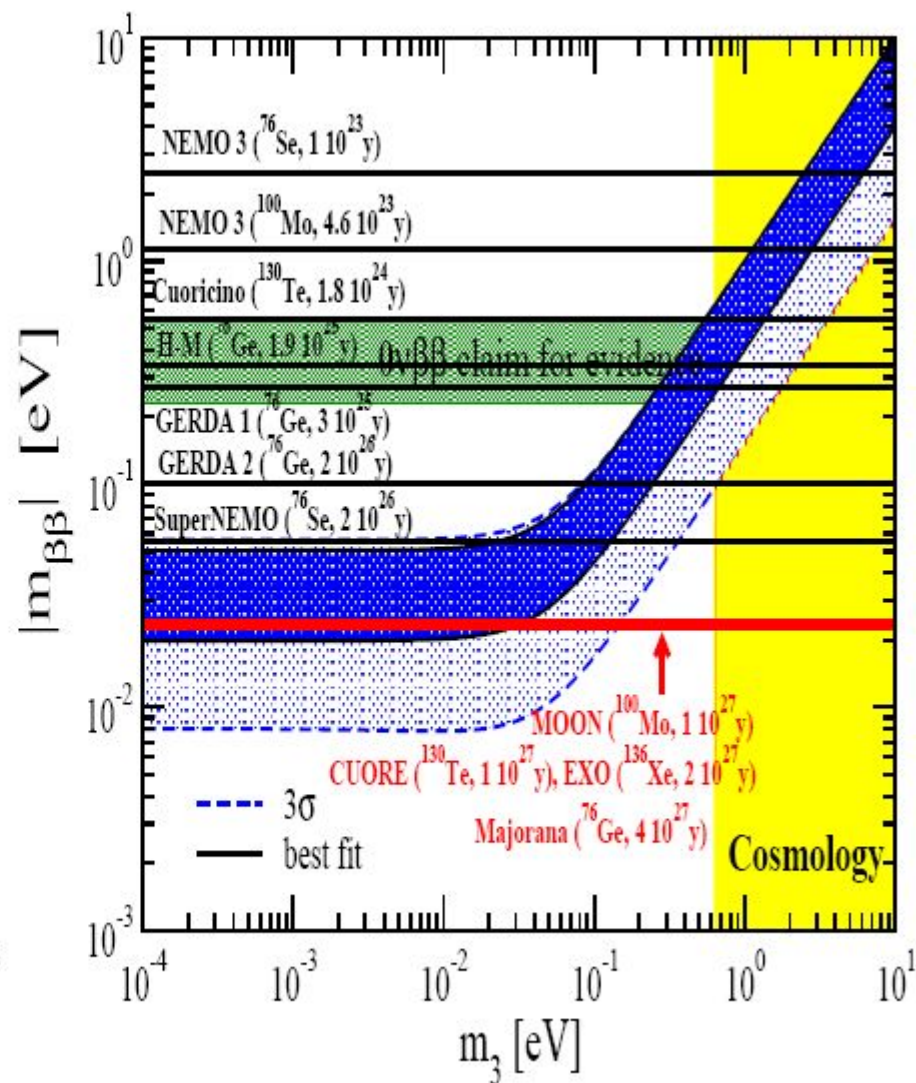
$\nu_e$  sector

- Bound from the CHOOZ experiment:  $\sin^2 2\theta_{13} < 0.1$
- studied also by MINOS & OPERA  $\rightarrow$  limited sensitivity
- next steps  $\rightarrow$  claim for factor 10 improvement :
  - T2K : appearance experiment
  - reactor experiments : disappearance experiment
- huge brainstorming for futur programs

# Mass hierarchy and $\beta\beta$ decay



Normal neutrino mass spectrum



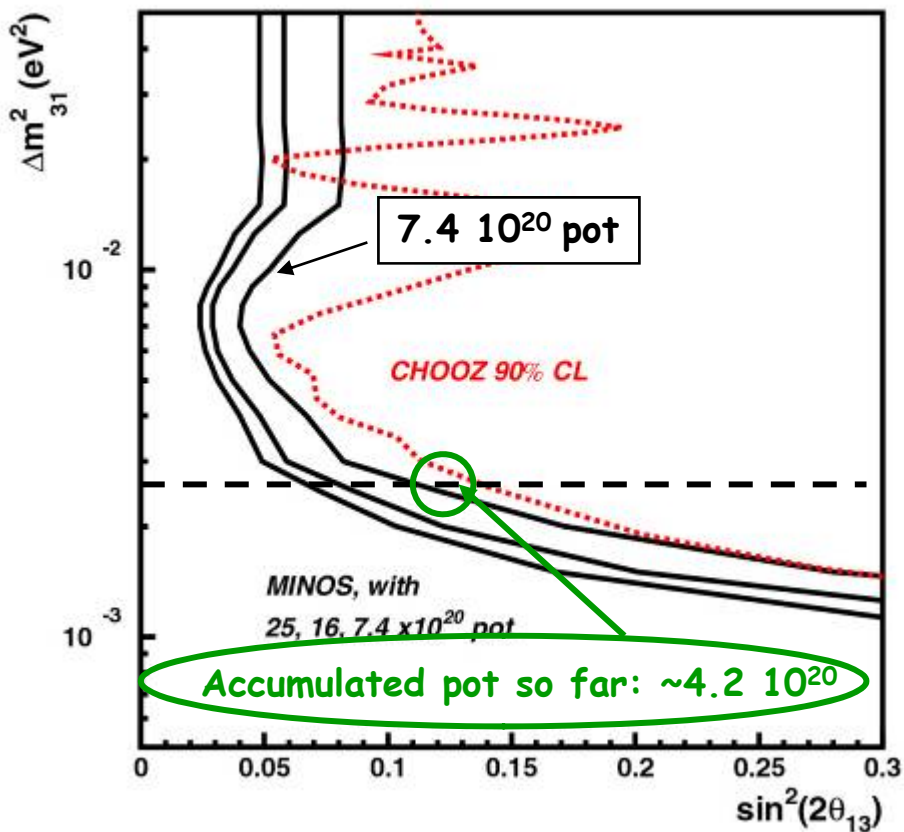
Inverted neutrino mass spectrum

# MINOS & OPERA sensitivity

## MINOS

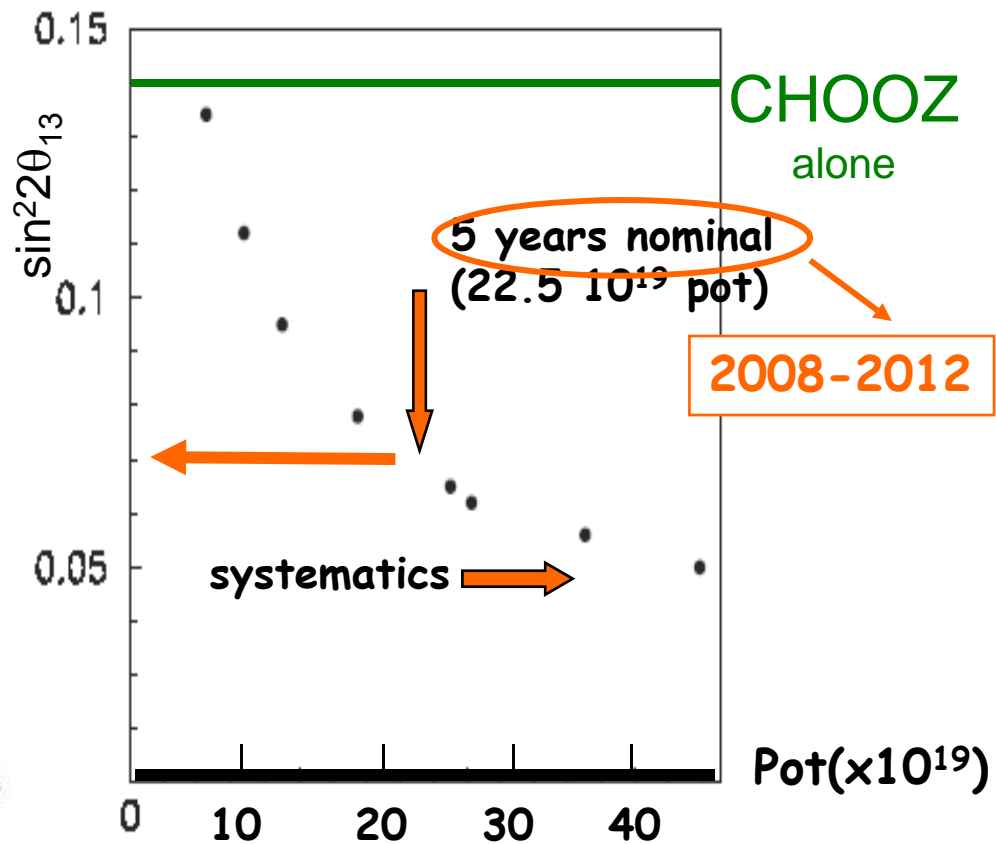
First results coming soon  
Analysis is improving

3  $\sigma$  Contours



## OPERA

Reduced sensitivity  
due to the 25% reduction  
of the target mass



# Reactor experiments

**Double Chooz:**  
(far detector construction started)

- Efficiencies included
- Systematics:
  - $\sigma_{\text{abs}} = 2.0\%$
  - $\sigma_{\text{rel}} = 0.6\%$
  - $\sigma_{\text{scl}} = 0.5\%$
  - $\sigma_{\text{shp}} = 2.0\%$
  - $\sigma_{\Delta m^2} = 20\%$
- **1% background subtraction error**

Assuming indirect measurements and a model for the background

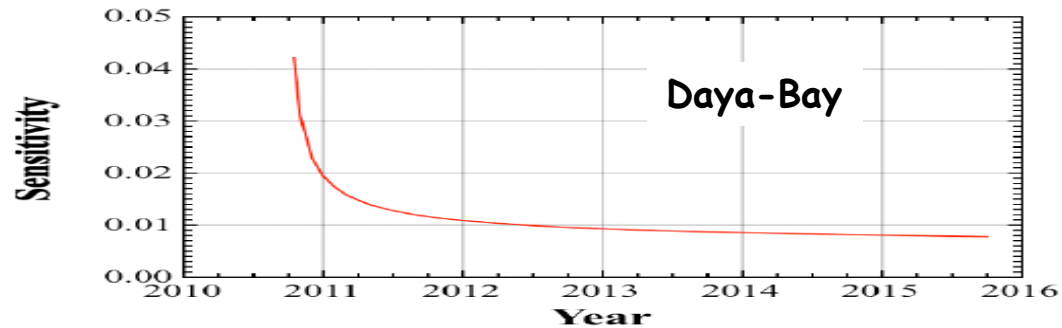
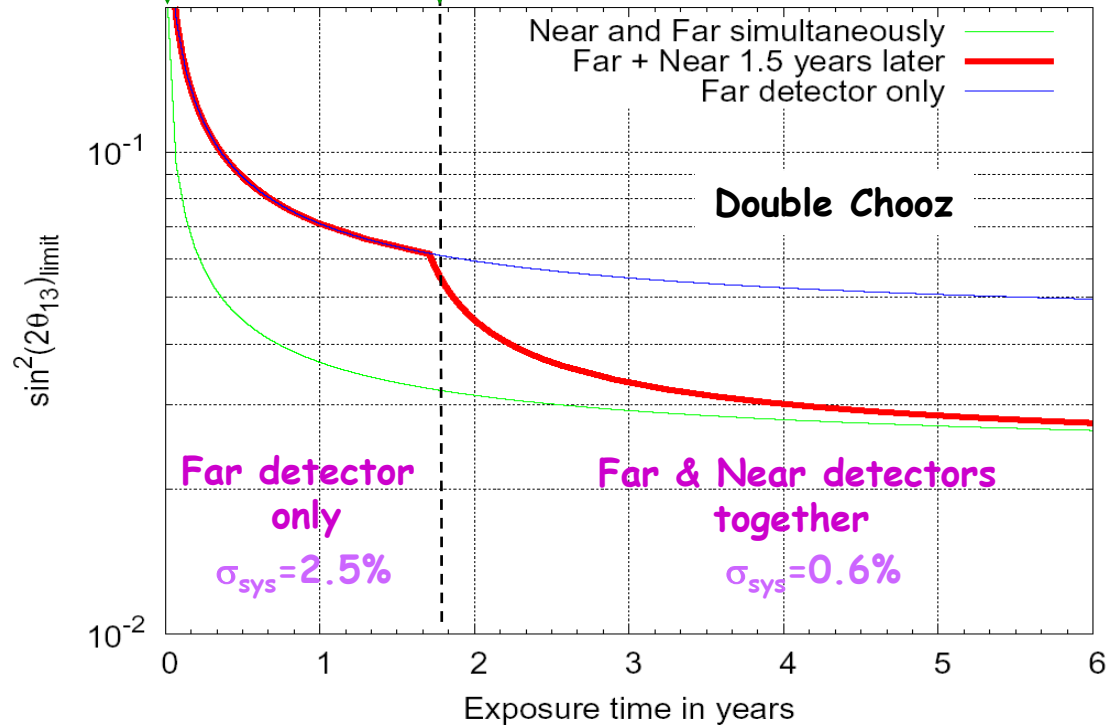
**Daya-Bay:**  
(excavation work starting)

- 0.2% background subtraction error

Assuming indirect measurements to prove the background negligible

August 2009

First half 2011

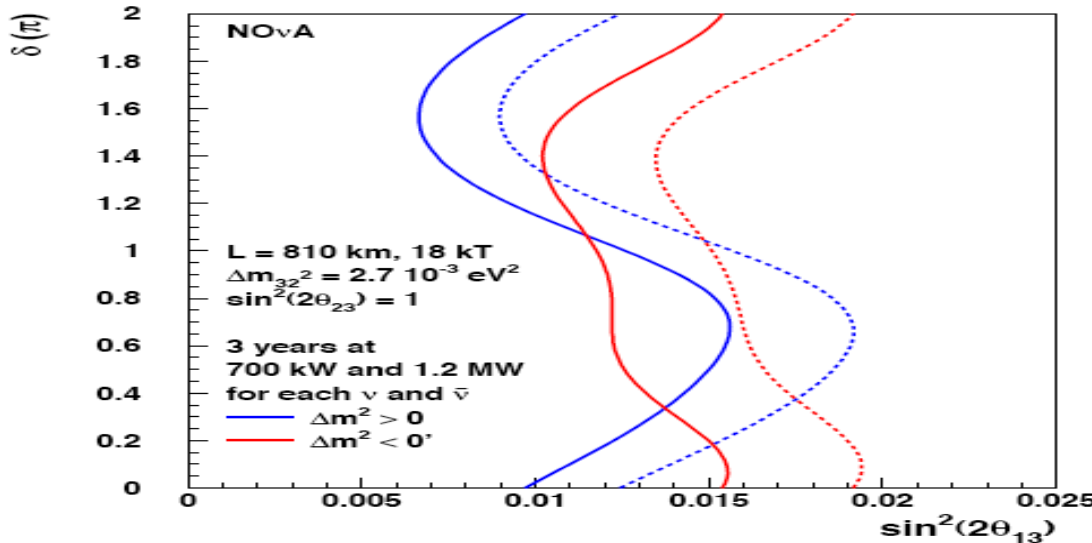
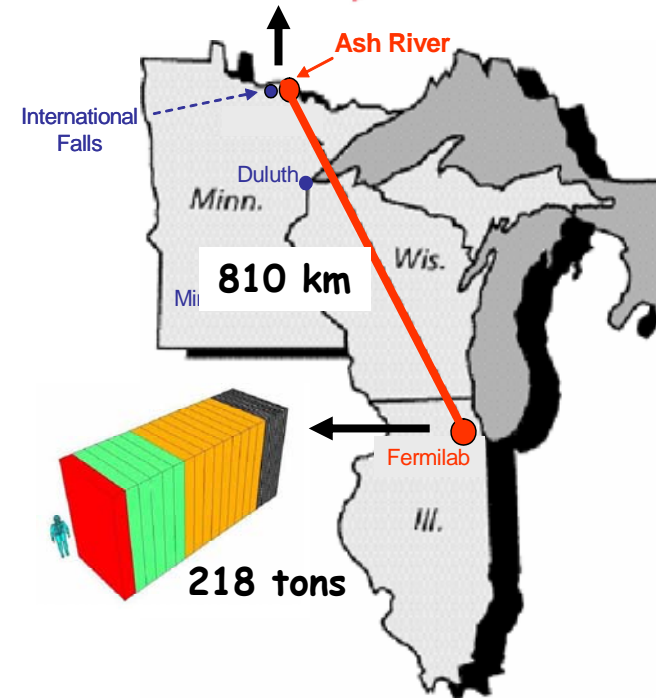
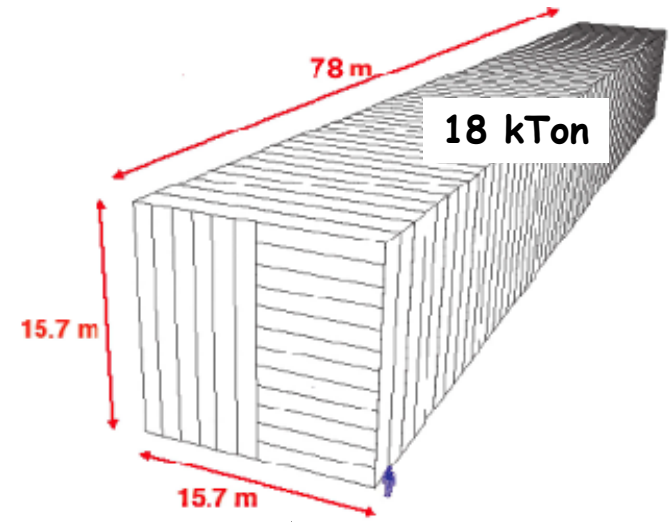


# Fermi Lab : NUMI upgrade

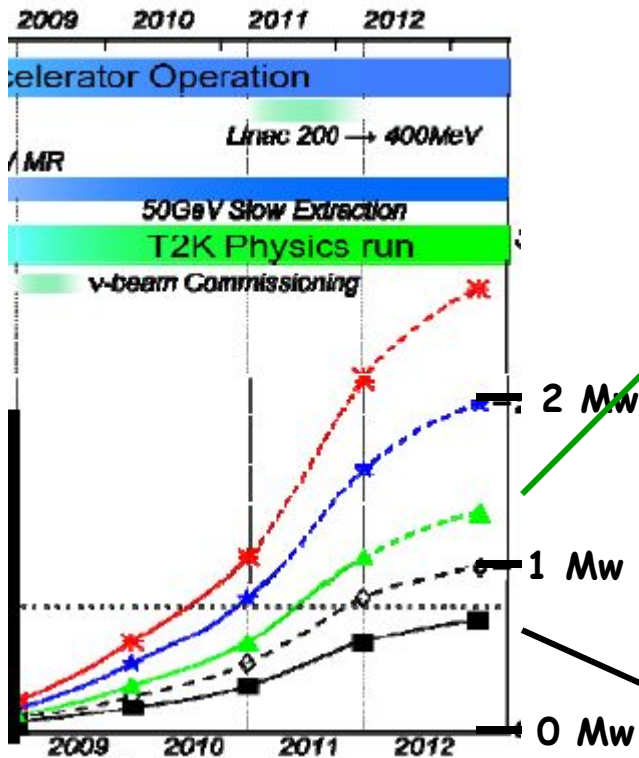
**NOvA** (beam + detectors :260 M\$)

- detectors : Totally active Lscint
- 14 mrad off axis
- NUMI beam line 320 kw  $\rightarrow$  700kw
- construction starting in 2010 ?
  - $\rightarrow$  critical budget in Fermi Lab:  
funding frozen for 1 year
- data taking in 2013 ?

$\rightarrow$  Focusing on Mass hierarchy measurement  
 $\rightarrow$  Strong complementarities with T2K  
*if some signal appears in the data*

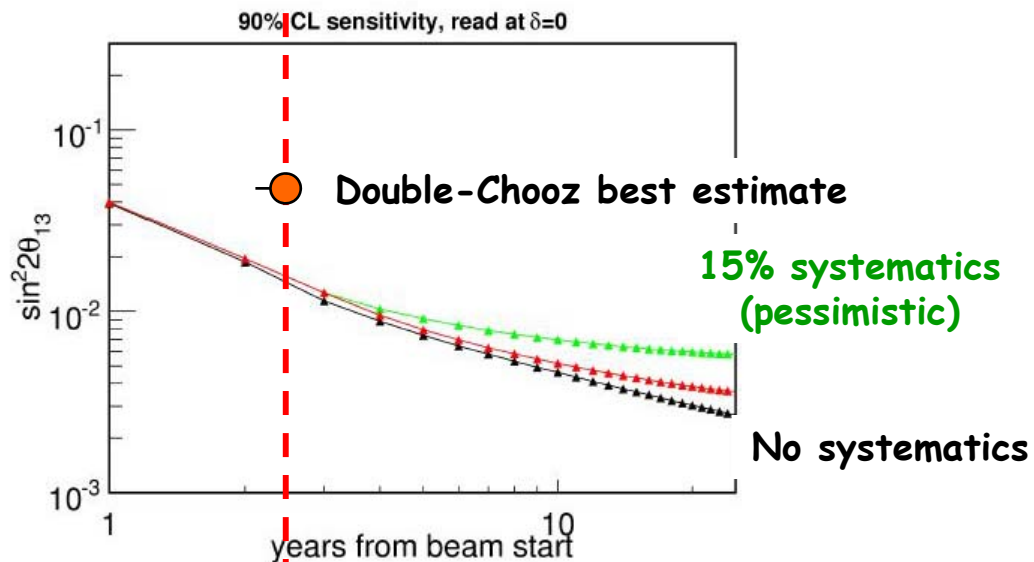


# T2K Phase 1 sensitivity

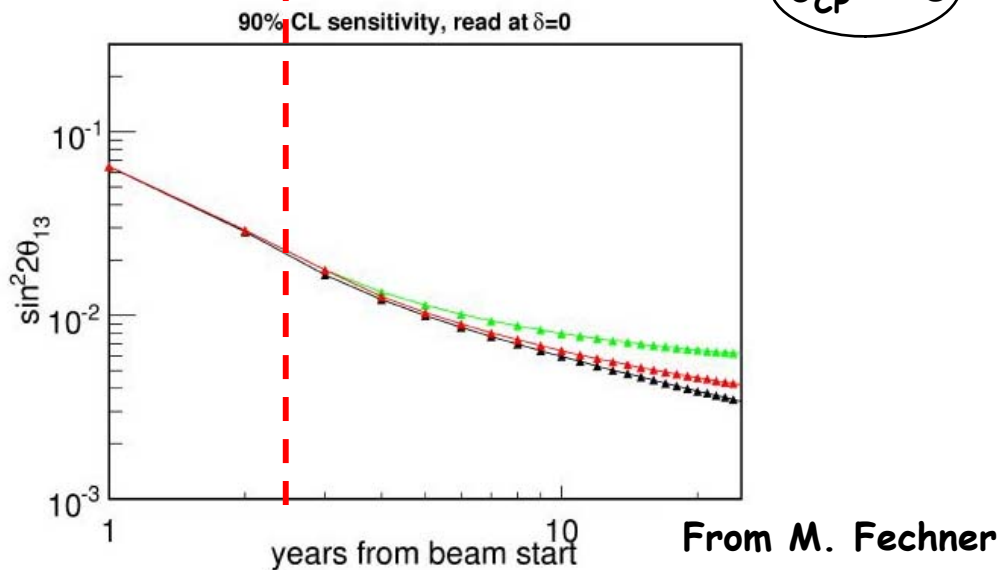


Down to  $10^{-2}$   
 the sensitivity of T2K  
 is mostly driven by statistics:  
 → Only driven by the integrated  
 flux delivered to SuperK

First half 2011



$$\delta_{CP} = 0$$



From M. Fechner



## Conclusions on T2K phase 1

By the end of 2011  
T2K can reach a sensitivity of  $1-2 \cdot 10^{-2}$   
for the  $\nu_{\mu} \rightarrow \nu_e$  transition in the atmospheric sector.

The integrated intensity delivered to the T2K program should be maximum during the first 3 years in order to accumulate the needed statistics and to guarantee the T2K discovery potential

Assuming a positive signal is observed then the next step is to determine the value of:

- $\theta_{13}$  angle
- the imaginary part : CPV
- the mass hierarchy

- increased statistics & anti-neutrino running
  - beam power
  - larger & improved detectors
- optimized baseline
- reducing the systematics

# Landscape for neutrino experiments after 2011

## Country driven programs:

- European situation
- Fermi Lab road map
- T2K phase 2

## International Scoping Studies for dedicated neutrino machines

- Beta beam
- Neutrino factory

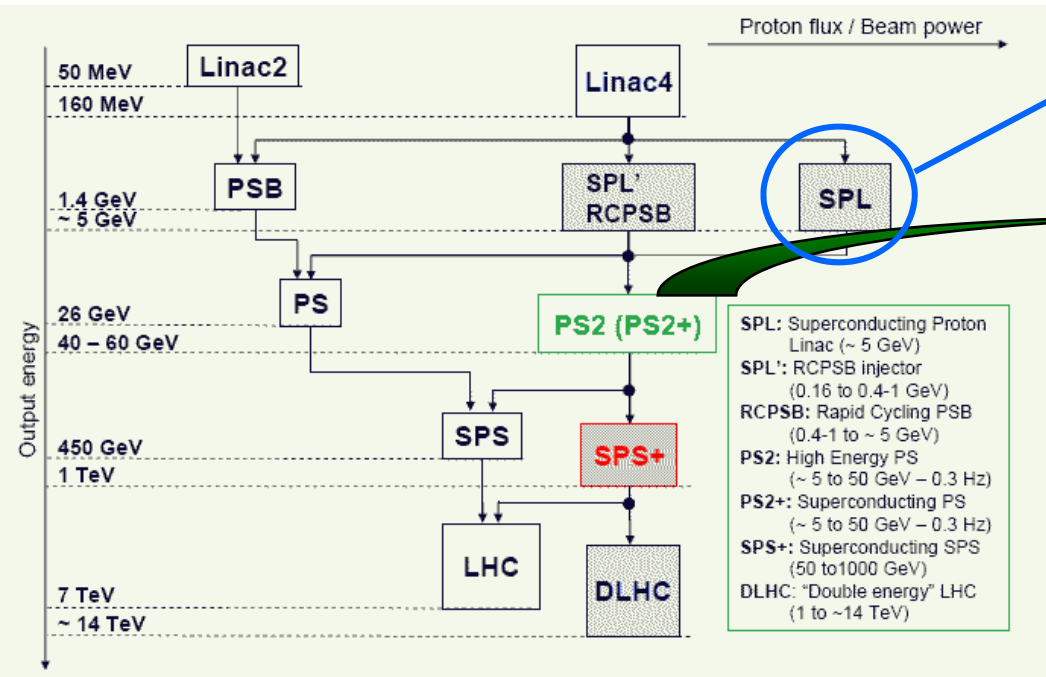
far beyond the scope of this talk  
(if we assume an effect observed in T2K phase 1)

# European Studies: a)SPS intensity increase

Not yet decided @CERN

Main motivation: LHC upgrade  
 → horizon > 2015

- injector for **EURISOL** →  $\beta$  beam
- Super Beam to Fréjus
- injector for  $\nu$ fact



[13] GLACIER : 100 Ktons LAr  
 [20] MODULAR : 30 Ktons LAr

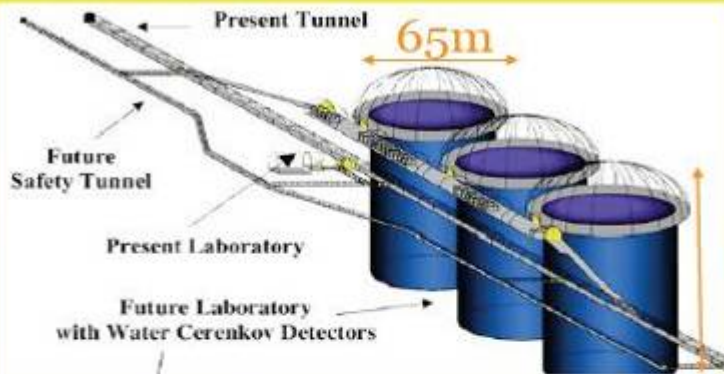
	CERN SpS			
	CNGS	+	1	2
Proton energy $E_p$	400 GeV			
$ppp(\times 10^{13})$	4.8	14	4.8	15
$T_c$ (s)	6	6	6	6
Efficiency	0.55	0.83	0.8	0.8
Running (d/y)	220	220	240	280
$N_{pot} / yr (\times 10^{19})$	7.6	33	12	43.3
Beam power (MW)	0.5	1.5	0.5	1.6
$E_p \times N_{pot}$ ( $\times 10^{22}$ GeV·pot/yr)	3	13.2	4.7	17.3
Relative increase		$\times 4$	$\times 1.5$	$\times 6$
Timescale	> 2008	>2016 ?		

## b) R&D for massive detectors funded by European Community

### Large Underground detectors considered in LAGUNA

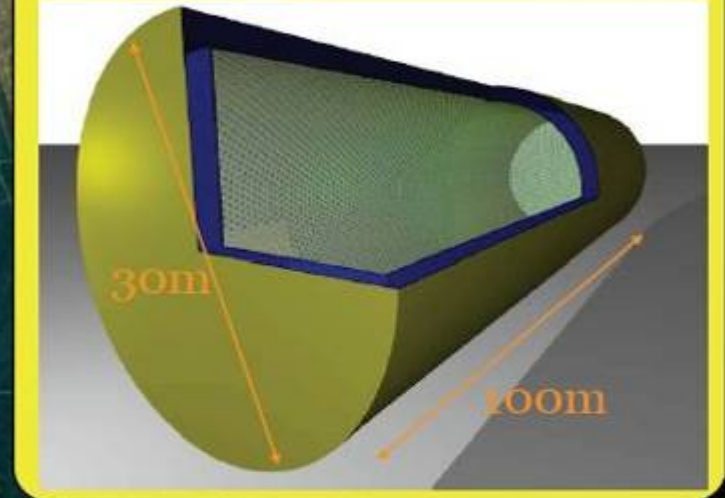
19

#### MEMPHYS-like



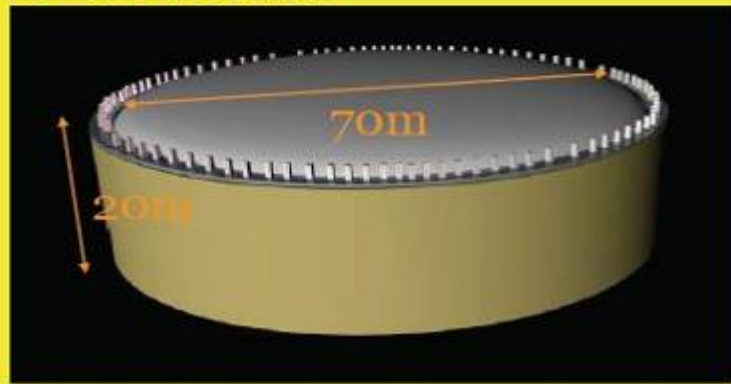
60m

#### LENA-like



Liquid Scintillator ( $\rightarrow$  50 kton)

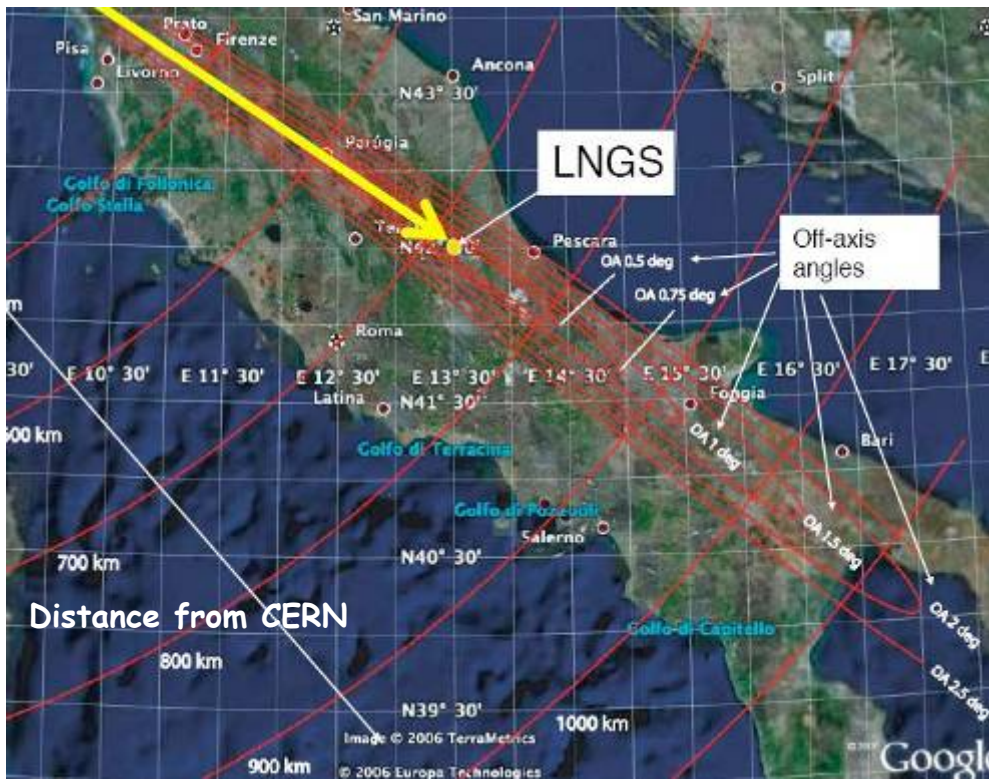
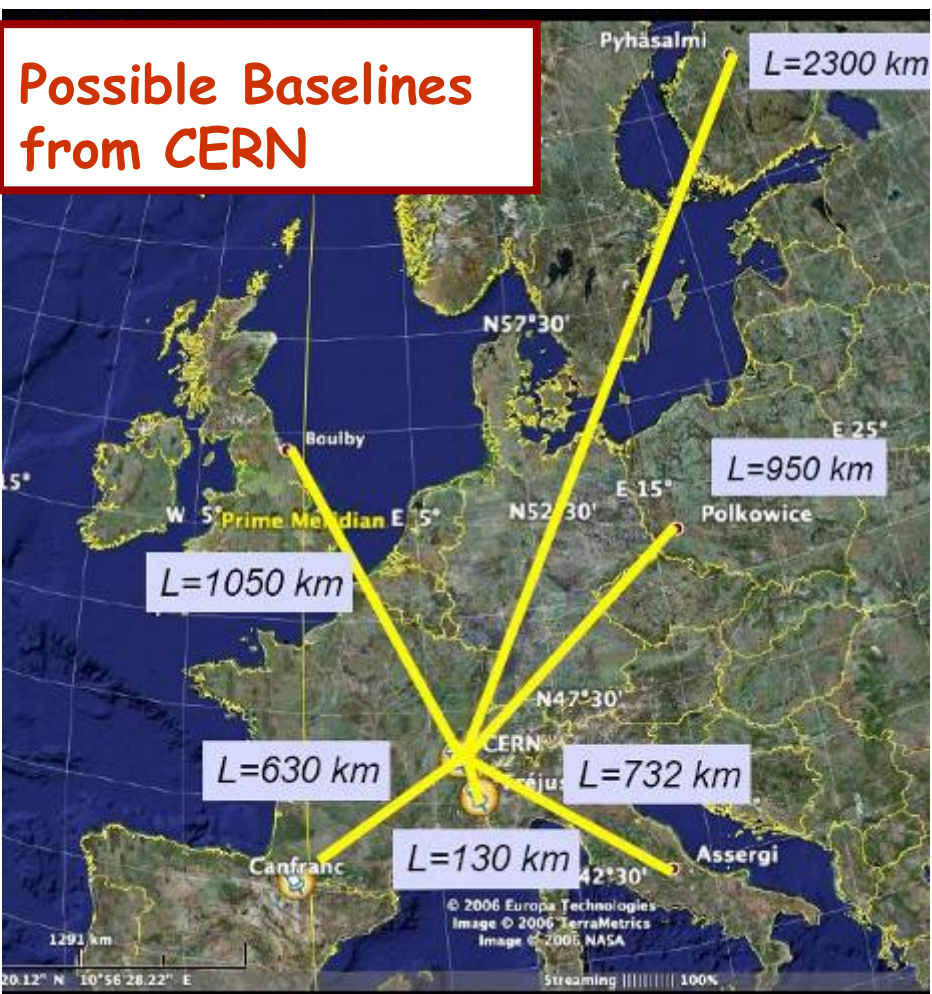
#### GLACIER-like



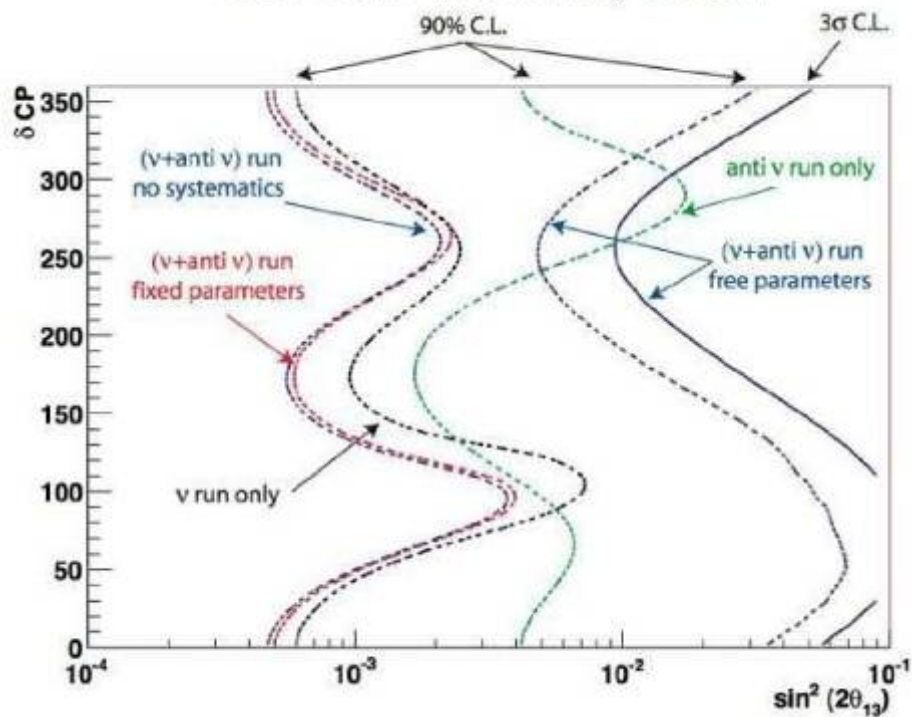
Liquid Argon ( $\approx$ 10  $\rightarrow$  100 kton)

photo: BOREXINO calibration

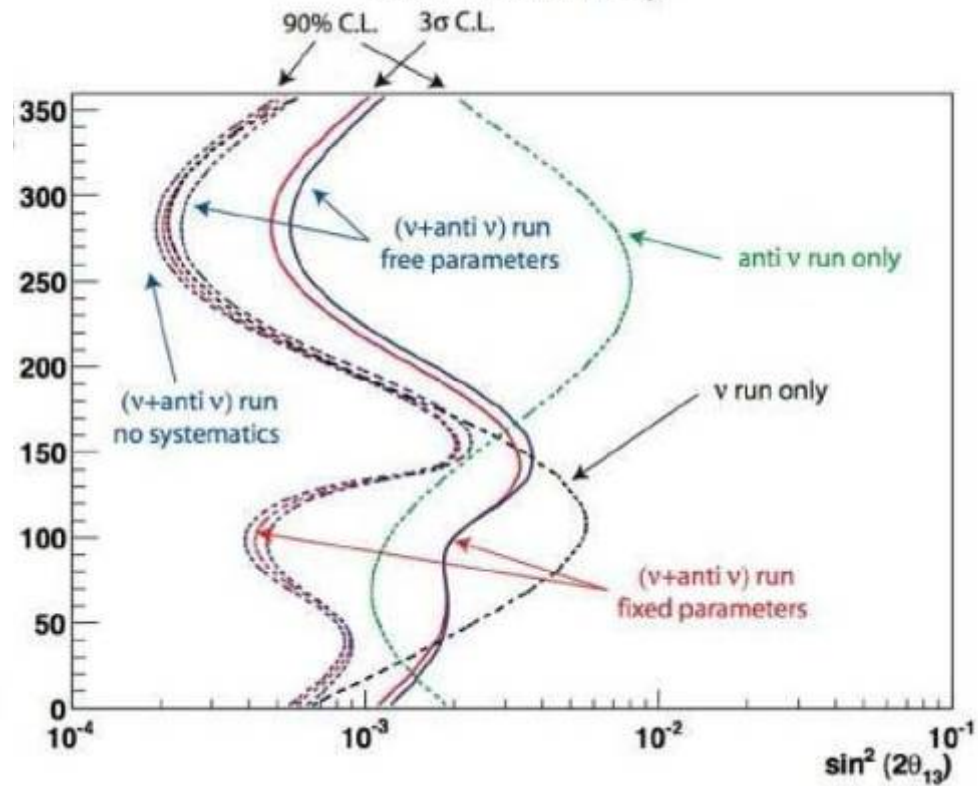
# Possible Baselines from CERN



CNGS - 850km - Mass hierarchy exclusion

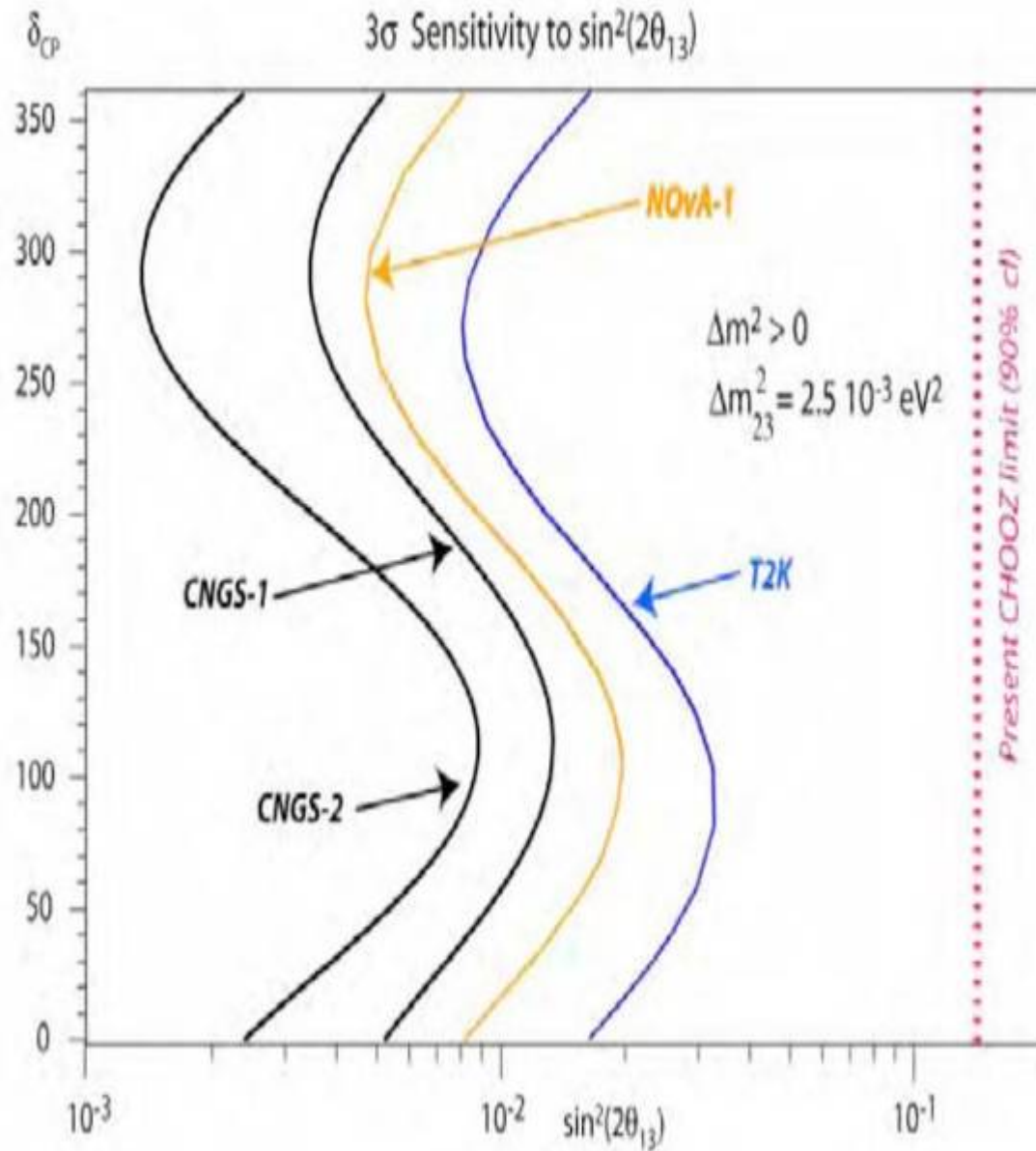


CNGS -  $\theta_{13}$  Discovery

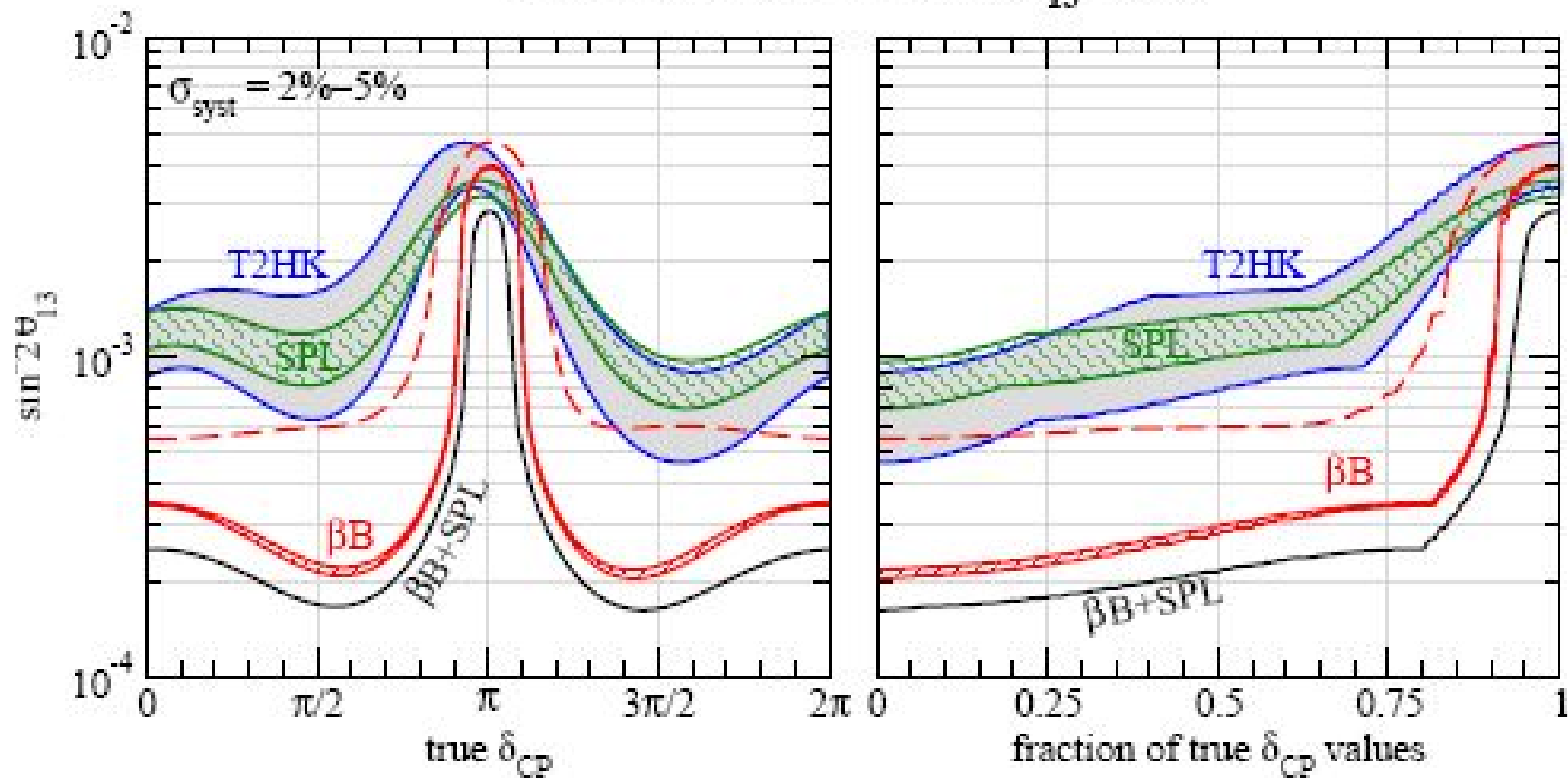




**T600 in the HallB @LNGS  
Should take data  
before the end of the year  
with the CNGS beam**

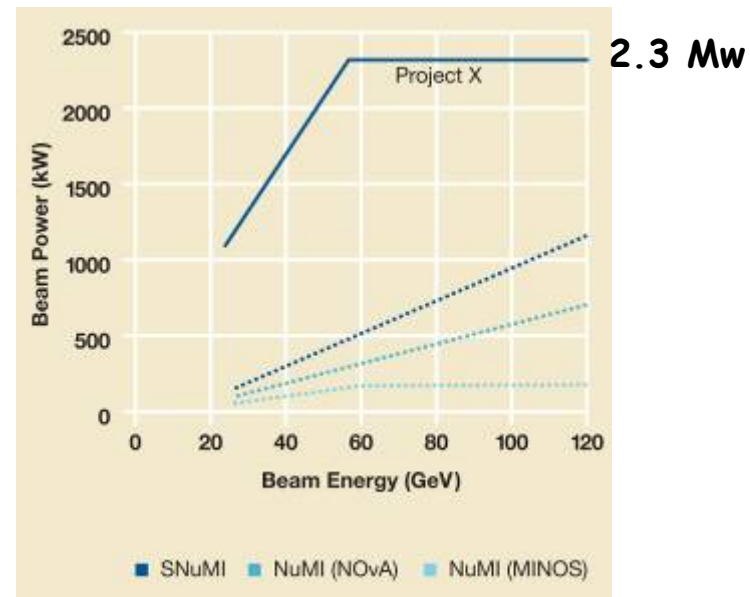
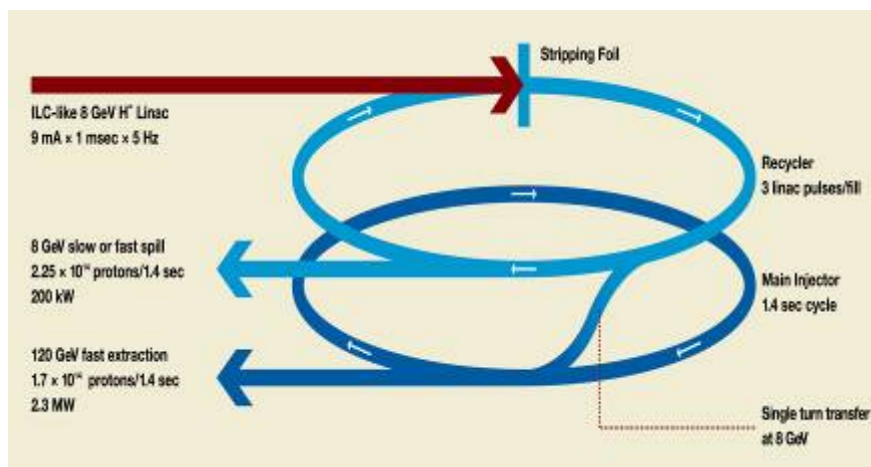


## Sensitivity to a non-zero $\theta_{13}$ at $3\sigma$



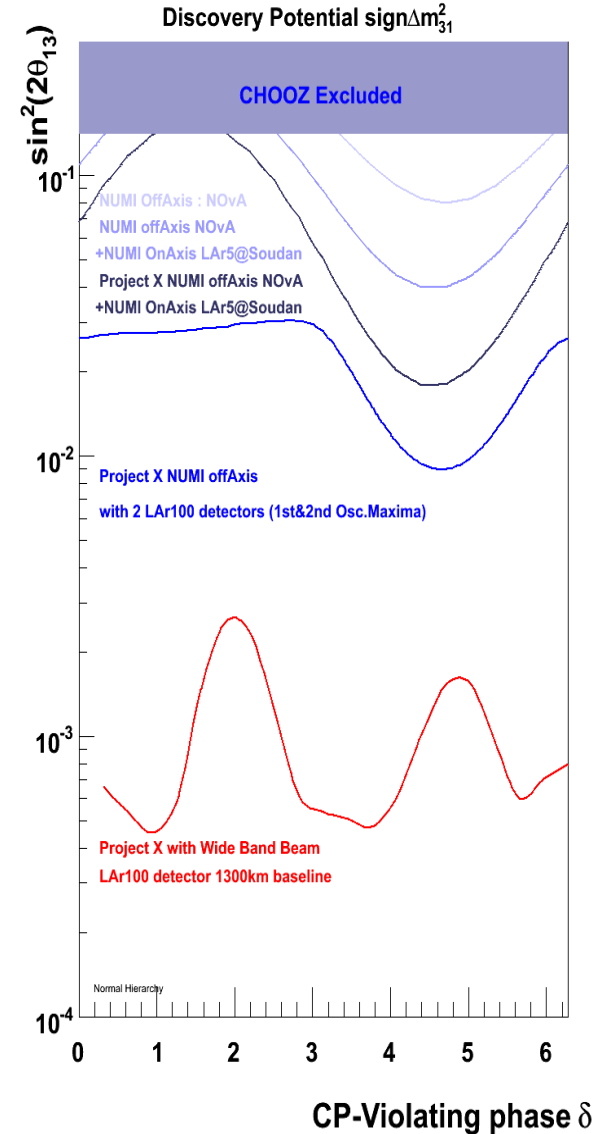
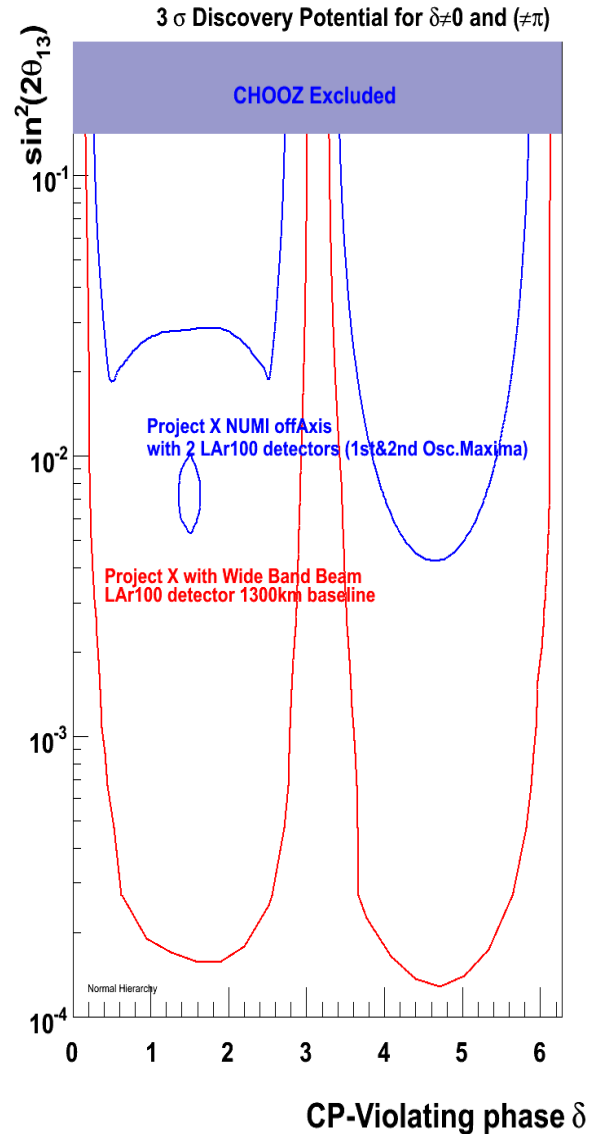
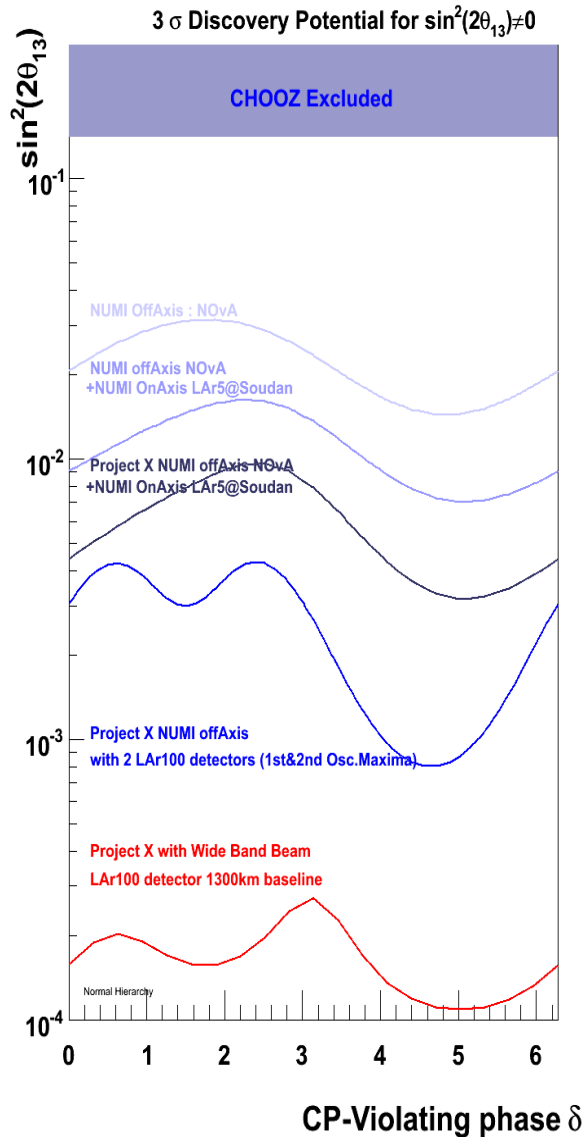


- 1) Breaking the intensity frontier  
→ multi Mw source for intense muon beam



- 2) Build NO<sub>v</sub>A
- 3) Replace MINOS by 5kTon LAr "on axis" → complementary to NO<sub>v</sub>A
- 4) Develop caverns/detectors for DUSEL  
with new beam-line from Project X it is the ultimate super-beam experiment (water or LAr)
- 5) If neutrino factory is needed - Project X is the ideal source  
as an example : [hep-ph-0709.3889](https://arxiv.org/abs/hep-ph-0709.3889) (A. Bross, M. Ellis, S. Geer) :  
A neutrino factory for both Large and Small  $\theta_{13}$

# Project X : sensitivity



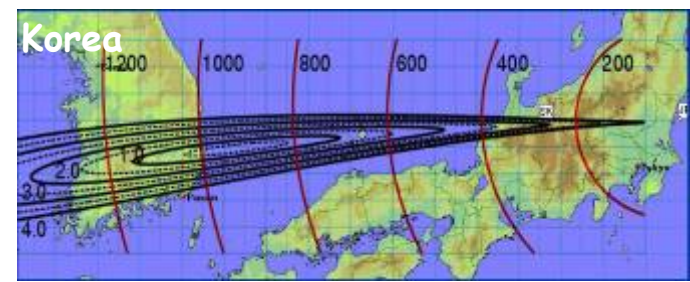
## Going to T2K phase 2

The goal of this workshop is to define the R&D needed for optimizing the strategy and be ready for a proposal in 2012

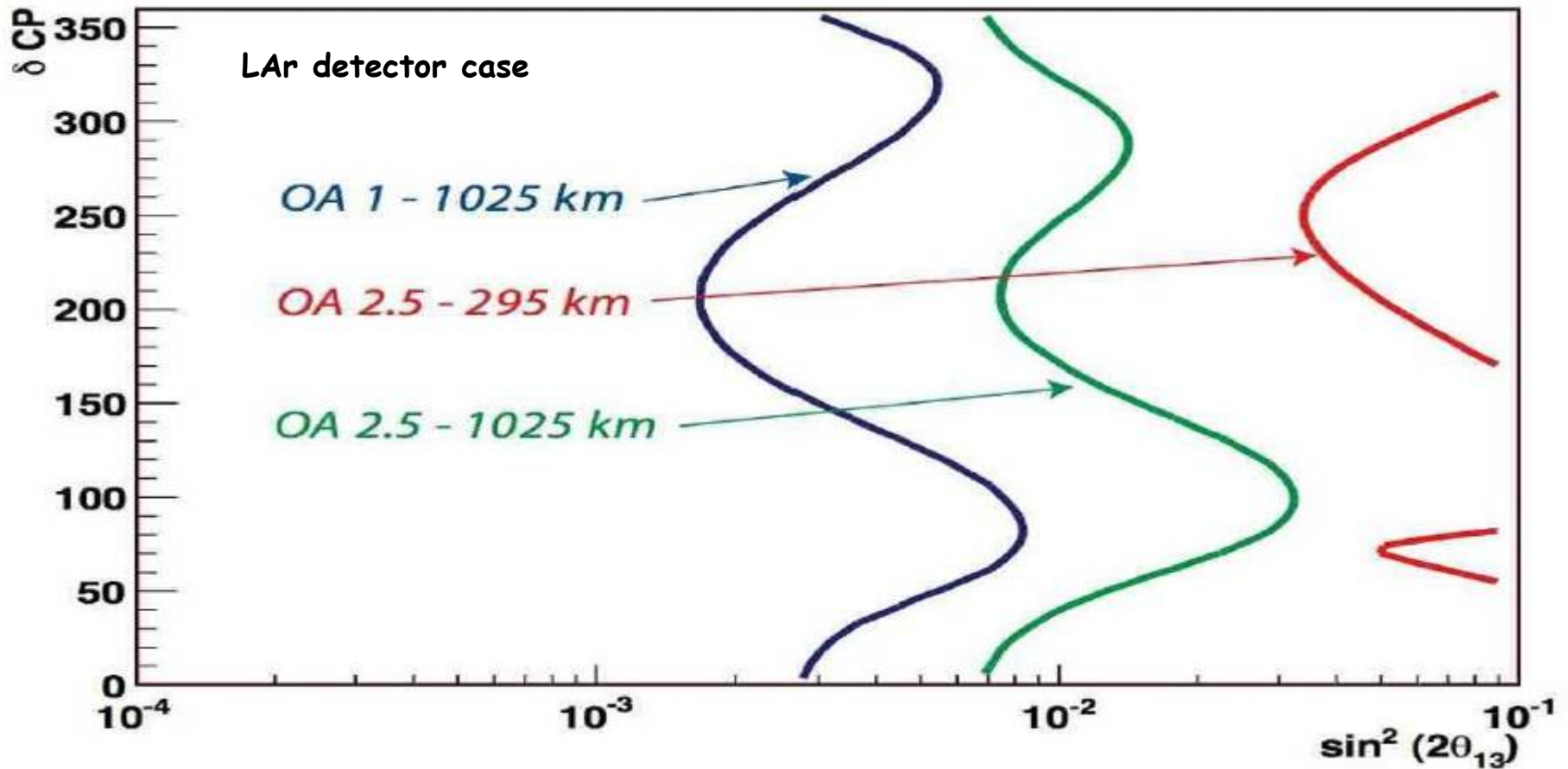
- 1) Intensity upgrade to 4Mw (~200 okuYen):
  - doubling the number of accelerated protons
  - repetition rate x2.5
  - proton energy : 40 GeV  $\rightarrow$  50 GeV

## 2) Optimizing the baseline

- going to the second maximum (T2KK)
  - ✓ More robust signal in the very far detector
  - ✓ Allows mass hierarchy sensitivity



Mass Hierarchy Determination - 4MW



## 2) Optimizing the Off Axis angle:

- Reducing the Off Axis angle
  - ✓ Wider energy distribution for the beam
  - ✓ Spectral information improve mass hierarchy sensitivity
  - ✓ Improve CPV sensitivity

Location	$\sin^2(2\theta_{13})$ = 0.002	$\nu_\mu$ CC + $\bar{\nu}_\mu$ CC no osc.	neutrino run				
			$\delta=0$	$90^\circ$	$270^\circ$	$180^\circ$	beam
J-PARC - 40 GeV/c protons - T2K optics - 4MW							
295 km 2.5 deg (0-5 GeV)	Matter (n.h.)	210970	274	39	393	158	4035 $\sqrt{B} = 64$
1025 km $\sim 1$ deg (0-5 GeV)	Matter (n.h.)	85900	226	138	389	300	776 $\sqrt{B} = 28$
1025 km 2.5 deg (0-5 GeV)	Matter (n.h.)	17475	94	60	126	92	334 $\sqrt{B} = 18$

Example with  
100 KTon LAr

$S/\sqrt{B}$  for  $\delta=0$

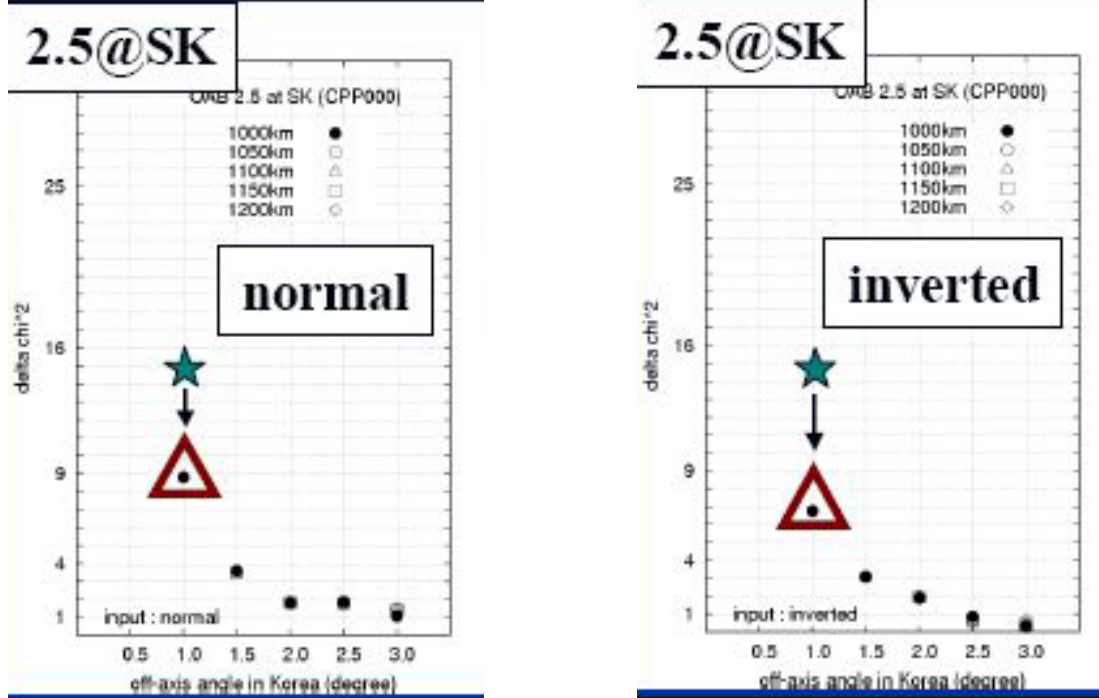
4.3

8.1

5.2

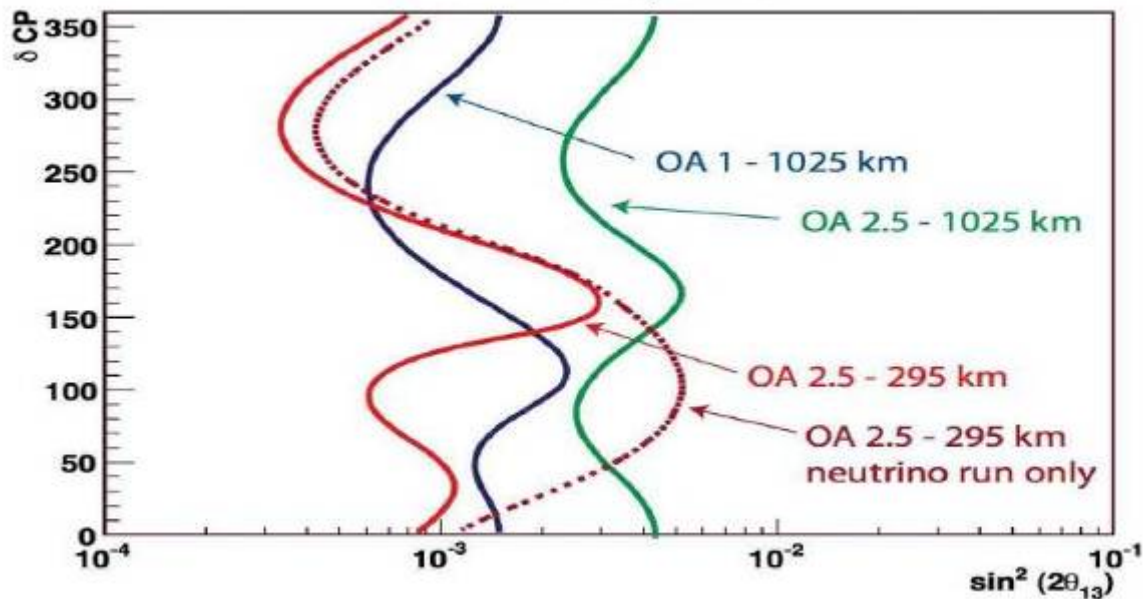
# Sensitivity

Water Cerenkov detector



$\theta_{13}$  Sensitivity - 4MW

LAr detector



## 4) Detector optimization

Not included in the topics of this talk

Water Cerenkov extrapolation from 50 Ktons → 1000 Ktons

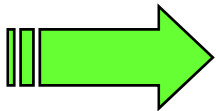
LAr TPC extrapolation from .3 Ktons → 200 Ktons

Technical  
Difficulties

To be weighted by

- Efficiencies
- NC Background rejection
- energy resolution
- accessing other physics fields like proton decay

Physics  
Performances



Do not forget systematics studies (beam & detector)  
with a near detector seeing the neutrino source as a point

# Conclusions

T2K phase1 is able to set quickly the scale  
for the study of the unknown part of the neutrino mixing matrix  
Unique discovery potential

In parallel with the running of T2K phase 1  
a vigorous & coherent R&D is mandatory  
for building a proposal for the detailed study of the effect  
(to be ready in 2012!)

This proposal should be strongly appealing  
in order to collect the largest worldwide neutrino community  
gathering so the needed resource