### Simulations for super raredecay experiments

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### MC Simulation

Estimate environment KL/neutron momentum, flux Design and evaluate beamline KL, core / halo neutron momeutum, flux Design detector Detector hit rate Requirements for detector inefficiency Quantitative evaluation of Signal/ Backgrounds Analysis stage

### Current E14 simulation status

Target simulation FLUKA/G3/G4 Beamline simulation FLUKA/G3/G4 Detector (G4 base framework) Kπνν signal K background • Kp2

- chKp3
- Ke3

Halo neutron background

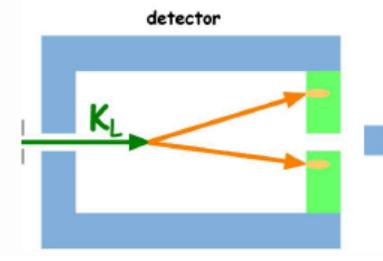
- CC02
- CV-π0

### $\pi^0 \nu \nu$ experiments

2 gammas from π<sup>0</sup> → Calorimeter

(In E14, π<sup>0</sup>→eeγ is not used due to small Br)

No other activities → Veto detectors
→ Two types of detectors



# MHz/cm<sup>2</sup> core neutron flux

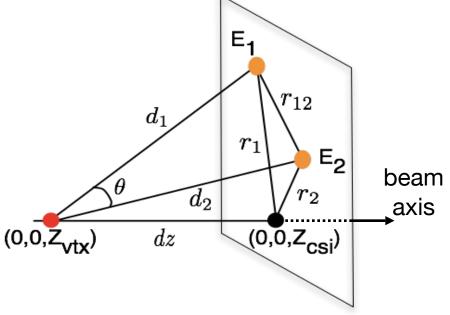
### E14 approach

#### Pencil beam method

Reconstruct p0 assuming its vertex is on the z axis

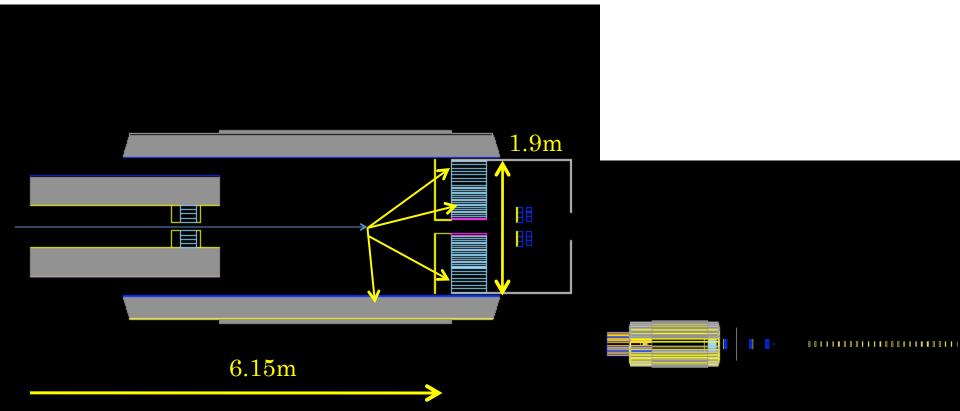
$$cos\theta = 1 - \frac{M_{\pi^0}^2}{2E_1E_2}$$
  
Calo:

- Energy leak
  - Shower leakage
  - Photonuclear effect
- Energy addition
  - Extra particles in  $\pi^0$  or  $\eta$  production by halo neutron



### E14 Detector

#### CsI : Calorimeter and Veto function



### Items to simulate

CsI as Calo

Gamma energy reconstruction

- Energy leak (Shower leakage and Photonuclear)
- Energy addition by extra particles  $\pi^0$  or  $\eta$  from neutron interaction

Cluster shape

CsI as Veto

Inefficiency (incl. fusion) down to 1 MeV

#### Other Veto detectors

Gamma inefficiency

- Inefficiency down to 1MeV
- Punch through, Sampling, Photonuclear effect Charged inefficiency
  - $\pi^{-} + A \rightarrow \pi^{0} + X$
  - $e^+ + e^- \rightarrow \gamma + \gamma$

Neutron interactions

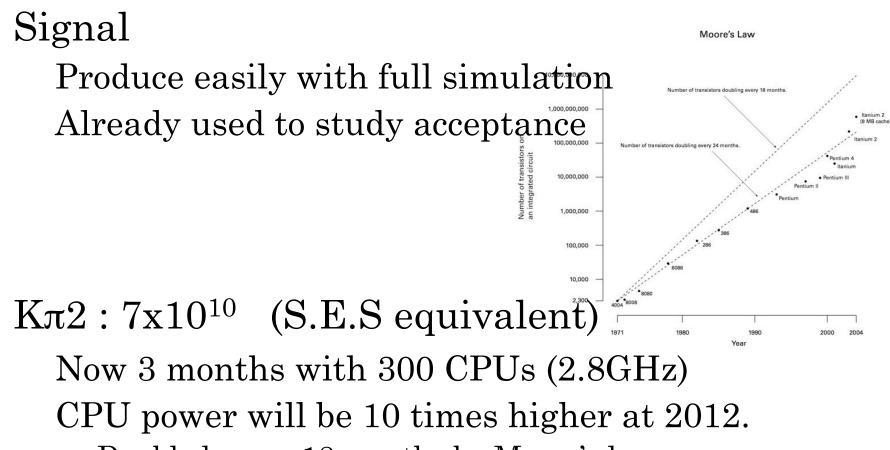
 $\pi^{\scriptscriptstyle 0}\!,\eta$  production

### E14 Simulation method

Full simulation

- Some speed-up methods
- Fast simulation
  - Inefficiency/fusion weight, smearing for CsI
- Recycling events of interest
- Fast-Full simulation
  - Use full simulation at CsI response and at halo-neutron interaction

### **Full Simulation**



• Doubled every 18 months by Moore's law.

### Speed up of MC production

| Time [sec] for                | Conditions  |
|-------------------------------|---|
| $10^4 K_L \rightarrow 2\pi^0$ |   |
| 120                           | FastSimLevel==5 for PhotonVetos   |
| 229                           | FastSimLevel==3 for PhotonVetos   |
| 1029                          | FastSimLevel==3 for BHPV  |
| 1099                          | FullShower  |
| 526                           | OnlineThresholdFastVeto   |
| 491                           | OnlineThresholdFastVeto, DetectorOrder  |
| 345                           | OnlineThresholdFastVeto, DetectorOrder, VtxFastTrigger                          |
|                               | $ ightarrow$ 9.2 months with 100 CPUs for $6.9	imes10^{10}~K_L ightarrow2\pi^0$ |

| ÷ | Stop at the surface of the volume.   |  |  |
|---|--|--|--|
| ÷ | Stop after the first step inside the volume.                               |  |  |
| ż | Stop 2ndary-creation from the processes,                                   |  |  |
|   | eBrem, annihil, and, conv.   |  |  |
| : | Kill the event immediately if Edep>Threshold.                              |  |  |
|   | CC00 0.05 GeV, CC01 0.05 GeV, CC02 0.05 GeV, CC03 0.05 GeV, CC04 0.05 GeV, |  |  |
|   | CC05 0.05 GeV, CC06 0.05 GeV, FBAR 0.05 GeV, CBAR 0.05 GeV, CV 0.002 GeV   |  |  |
| 2 | Decide from which detector shower simulation start.                        |  |  |
|   | BHCV CV BCV CC04 CC05 CC06 BHPV CC00 CC01 CC02 FBAR CBAR CSI               |  |  |
| 2 | Trigger with the end z-position of the primary.                            |  |  |
|   | /GsimStackingAction/triggerPrimaryEndZ -10*m 10*m                          |  |  |
|   |  |  |  |

### **Fast Simulation**

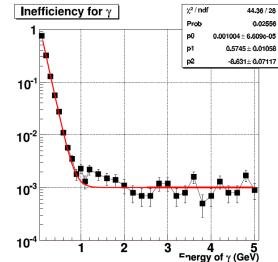
Stop particles on the surface of detectors Use inefficiency, fusion weights

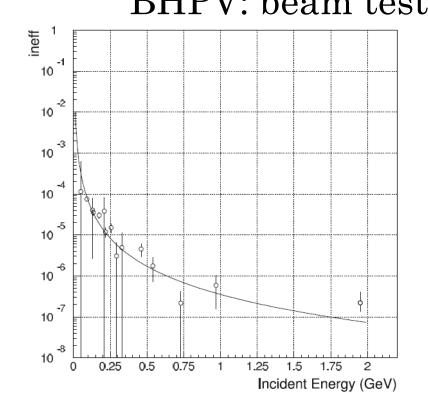
- It enables to evaluate S/N with smaller MC statistics
  - Rare event topologies can be treated as event weights.
  - BG evaluation with  ${\sim}10^7$  events / 1day for K backgrounds.

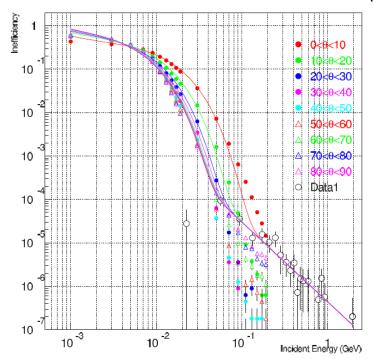
|           |                                    |               |                       | acceptance loss |
|-----------|------------------------------------|---------------|-----------------------|-----------------|
|           |                                    | standard cuts | CsI cluster shape cut | (50%)           |
| Signal    | $K_L \to \pi^0 \nu \overline{\nu}$ | $6.0 \pm 0.1$ | $5.4 \pm 0.1$         | $2.70\pm0.05$   |
| $K_L$ BG  | $K_L \to \pi^0 \pi^0$              | $3.7\pm0.2$   | $3.3\pm0.2$           | $1.7\pm0.1$     |
|           | $K_L \to \pi^+ \pi^- \pi^0$        | $0.18\pm0.08$ | $0.16\pm0.07$         | $0.08\pm0.04$   |
|           | $K_L \to \pi^- e^+ \nu_e$          | $0.13\pm0.01$ | $0.03\pm0.003$        | $0.02\pm0.001$  |
| halo n BG | CV                                 |               |                       | 0.08            |
|           | $\eta$                             | 8.1           | 0.6                   | 0.3             |

## Gamma Inefficiency Function

CsI, Veto, BHPV Photonuclear from Exp.Data Veto : sampling calo. MC simu. BHPV: beam test and MC





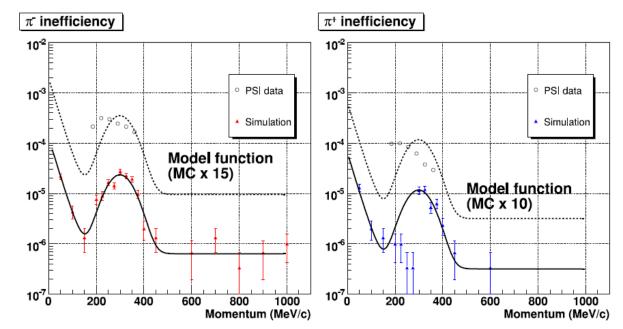


### **Charged Inefficiency Function**

CV: PSI measurements with a photon calorimeter behind.

• Inconsistency due to the small size of a calorimeter used in the PSI measurements.

BHCV : 0.5% by masking effect due to high counting rate



### Recycle events of interest

Multi-stage MC production for halo-neutron backgrounds and Ke3 (conversion type) For example, CC02 backgrounds

Halo n $\rightarrow$ Hit at CC02 Produce  $\pi 0$  Veto $\rightarrow 2\gamma$  in CsI $\rightarrow \pi 0$  reconstruction

CC02 hitCC02 x pi0pord x veto 50%  $10^{-3}$   $10^{-5} \rightarrow 10^{8}$  gain CV-eta hitCV x eta prod x 2ybranch x veto 10%  $10^{-6}$  0.4  $1\% \rightarrow 10^9$  gain CV- $\pi$ 0: hitCV x  $\pi$ 0prod x veto 10%  $10^{-4}$  1%  $\rightarrow$   $10^{7}$  gain Ke3 : hitCV x  $\pi^{-}$  ineff. x e+ ineff. x reconstruction  $4\% \ge 10^{-4} \ge 10^{-4} \ge 30\% \rightarrow 10^{10}$  gain  $\rightarrow$  2-3 days to treat 10 times E14 halo neutrons

### **Fast-Full Simulation**

Use full simulation at CsI response and at haloneutron interaction

Merit:

Shower can be treated easily.

- Cluster shape cut
- Inefficiency v.s. Collateral cluster
- Backsplash loss (They are not independent.)

Extra particles at neutron interactions

Demerit:

Weight can't used  $\rightarrow$  Need large statistics to treat rate event topologies.

Anyway, we are getting more realistic information around cluster shape cut

### Future plan

Speed up

Shower library for inefficient gammas in  $CsI \rightarrow It$  enables weight method to treat rare event topologies.

Complete detector design and construct detector. Calibration method for inefficiency function in data acquisition.

Step2: 100 times higher statistics

Still affordable if these methods are validated.

→~1 month for generation incl. Moor's law.